

# Frequency evaluation of Maser 1401104 by IT-Yb1 for the period MJD 59669 to 59679

During the period MJD 59669 – 59679 (31 March 2022–10 April 2022) INRIM evaluated the frequency of the hydrogen maser IT-HM4 (BIPM code 1401104) using the Yb optical lattice frequency standard IT-Yb1 and an optical frequency comb. The evaluation is based on the recently published **CCTF2021** recommended frequency for <sup>171</sup>Yb as a secondary representation of the second,  $f(^{171}Yb) = 518\,295\,836\,590\,863.63\,\text{Hz}$  with a relative standard uncertainty of  $u_{\text{Srep}} = 1.9 \times 10^{-16}$  [1, 2]. The results of the evaluation are summarized in Tab. 1. Details of IT-Yb1 operation and uncertainty budget are given in Refs. [3, 4] and summarized below.

## 1 Frequency measurement

The clock laser of IT-Yb1 is stabilized on an ultrastable cavity and probes <sup>171</sup>Yb atoms trapped in an optical lattice at the magic frequency. A digital control loop acting on an acousto-optic modulator keeps the clock laser frequency in resonance with the atoms. The cavity-stabilized laser is sent to a fibre frequency comb referenced to IT-HM4. The frequency ratio between the <sup>171</sup>Yb transition and IT-HM4 is calculated from the comb measurements and the corrections used for steering the acousto-optic modulator.

Table 1: Final evaluation using IT-Yb1.							
Period of es- timation	y(IT-HM4)/ITYb1	$u_{\mathrm{A}}$	$u_{\rm B}$	$u_{\rm A/lab}$	$u_{\rm B/lab}$	$u_{\rm Srep}$	Uptime
	$/10^{-16}$	$/10^{-16}$	$/10^{-16}$	$/10^{-16}$	$/10^{-16}$	$/10^{-16}$	
59669 - 59679	-437.6	0.0	0.2	1.3	0.2	1.9	45.6%

Effect	Rel. Shift/ $10^{-17}$	Rel. Unc./ $10^{-17}$
Density shift	-1.2	0.3
Lattice shift	5.0	1.6
Zeeman shift	-3.13	0.03
Blackbody radiation shift (room)	-235.2	1.1
Blackbody radiation shift (oven)	-1.3	0.6
Static Stark shift	-1.5	0.9
Probe light shift	0.04	0.03
Background gas shift	-0.5	0.2
Servo error	0.0	0.3
Other shifts	0.0	0.1
Grav. redshift (static)	2599.5	0.3
Grav. redshift (tides)	0.0	0.2
Total	2361.7	2.3

Table 2: Uncertainty budget for IT-Yb1 for the reported period.

Table 3: Uncertainty budget for the link between IT-Yb1 and IT-HM4 for the reported period.

Effect	Uncertainty/ $10^{-16}$
Comb statistic	0.1
Extrapolation (dead time)	1.3
Extrapolation (drift)	0.0
Total $u_{\rm A/lab}$	1.3
Optical/microwave comp. (type B)	0.2
Total $u_{\rm B/lab}$	0.2

# 2 IT-Yb1 evaluation

The uncertainty  $u_{\rm A}$  is the statistical contribution from the instability of IT-Yb1. The uncertainty  $u_{\rm B}$  is the systematic uncertainty of IT-Yb1 [4]. The systematic frequency shift and uncertainty budget of IT-Yb1 for the reported period are given in Tab. 2. IT-Yb1 now operates with a vertical optical lattice and the lattice light shift calculations have been updated following Ref. [5]. The table includes the gravitational redshift relative to the conventional potential  $W_0 = 62\,636\,856.0\,\mathrm{m}^2\mathrm{s}^{-2}$  [4].

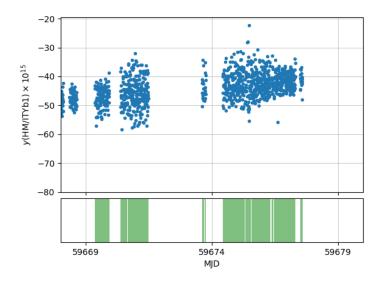


Figure 1: Fractional frequency deviation y(HM1401104/ITYb1) measured in the period MJD 59669 - 59679. Green shaded regions in the bottom plot represent the uptime of IT-Yb1.

## 3 Link evaluation

The uncertainty  $u_{l/lab}$  is due to the link between IT-Yb1 and IT-HM4, including the optical to microwave comparison at the comb. Table 3 summarizes the contributions to this uncertainty.

The comparison uncertainty between optical and microwave signals at the comb has been evaluated from comparison with a second optical frequency comb.

IT-Yb1 and the comb were operated for  $394\,330\,\mathrm{s}$  (uptime 46% of the evaluation period). The data collected and the distribution of the uptimes of IT-Yb1 are shown in Fig. 1. Extrapolation using the maser as a flywheel is needed given the intermittent operation of IT-Yb1. Its evaluation is separated in an uncertainty from dead times and a correction for the maser drift. The maser drift of  $7.6(4) \times 10^{-16}$  /d has been calculated from IT-Yb1 data collected in the period. The contribution from dead times has been evaluated following the approach in Ref. [6]. For this measurement we considered the IT-HM4 noise to be a power-law model described by the Allan deviation: white phase noise  $3 \times 10^{-13} (\tau/\mathrm{s})^{-1}$ ; white frequency noise  $4 \times 10^{-14} (\tau/\mathrm{s})^{-1/2}$ ; flicker frequency noise  $3 \times 10^{-16}$ ; random walk frequency noise  $2 \times 10^{-19} (\tau/\mathrm{s})^{1/2}$ .

#### Contributors

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